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Validation of the McClear clear-sky model in desert conditions with three stations in Israel

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Abstract. The new McClear clear-sky model, a fast model based on a radiative transfer solver, exploits the atmospheric properties provided by the EU-funded Copernicus Atmosphere Monitoring Service (CAMS) to estimate the solar direct and global irradiances received at ground level in cloud-free conditions at any place any time. The work presented here focuses on desert conditions and compares the McClear irradiances to coincident 1 min measurements made in clear-sky conditions at three stations in Israel which are distant from less than 100 km. The bias for global irradiance is comprised between 2 and 32 W m⁻², i.e. between 0 and 4 % of the mean observed irradiance (approximately 830 W m⁻²). The RMSE ranges from 30 to 41 W m⁻² (4 %) and the squared correlation coefficient is greater than 0.976. The bias for the direct irradiance at normal incidence (DNI) is comprised between -68 and +13 W m⁻², i.e. between -8 and 2 % of the mean observed DNI (approximately 840 W m⁻²). The RMSE ranges from 53 (7 %) to 83 W m⁻² (10 %). The squared correlation coefficient is close to 0.6. The performances are similar for the three sites for the global irradiance and for the DNI to a lesser extent, demonstrating the robustness of the McClear model combined with CAMS products. These results are discussed in the light of those obtained by McClear for other desert areas in Egypt and United Arab Emirates.

1 Introduction

The downwelling solar irradiance observed at ground level on horizontal surfaces and integrated over the whole spectrum (total irradiance) is called surface solar irradiance (SSI). It is the sum of the direct irradiance, from the direction of the sun, and the diffuse, from the rest of the sky vault, and is also called the global irradiance. The SSI is an essential climate variable as established by the Global Climate Observing System in August 2010 (GCOS, 2016). Knowledge of the SSI and its geographical distribution is of prime importance for numerous domains where SSI plays a major role as e.g. weather, climate, biomass, and energy.

A model estimating the SSI under clear sky or cloud-free conditions is called a clear-sky model. Oumbe et al. (2014) have demonstrated that computations of the SSI from satellite images can be approximated by the product of the clearsky SSI and a modification factor due to cloud properties and ground albedo only. Changes in clear-atmosphere properties have negligible effect on this modification factor so that both terms can be calculated independently. These results are important in the view of an operational system as it permits separating the whole processing into two distinct and independent models, whose input variable types and resolutions may be different. This enforces the importance of the availability of an accurate and easy-to-operate model for the assessment of the clear-sky SSI.

The McClear model (Lefèvre et al., 2013) is such a model. It has been designed to benefit from the recent advances on atmosphere composition made in MACC projects (Monitoring Atmosphere Composition and Climate). The latter were preparing the operational provision of global aerosol properties analyses and forecasts together with physically consistent total column content in water vapour and ozone available every 3 h (Benedetti et al., 2009; Kaiser et al., 2012; Peuch et al., 2009). Such information had not been available so far from any operational numerical weather prediction centre. Since 1 January 2016, the McClear model and its inputs are part of the operational services delivered by the Copernicus Atmosphere Monitoring Service (CAMS) operated by ECMWF on behalf of the European Commission.

Station	Latitude (positive North, ISO 19115)	Longitude (positive East, ISO 19115)	Elevation a.s.l. (m)
Beer Sheva (BEE)	31.25	34.8	195
Sede Boqer (SBO)	30.905	34.782	500
Yotvata (YOT)	29.879	35.065	66

Table 1. Geographical coordinates of the three stations. Period is 2006–2011. All data are coincident. Number of samples is 19 849 in *G*, *B* and *D*.

The CAMS McClear service is available as an interoperable Web processing service (WPS), i.e. an application that can be invoked via the Web and that obeys the OGC (Open Geospatial Consortium) standard for interoperability (Percivall et al., 2011). This service delivers estimates of the global SSI and its direct and diffuse components on horizontal surface as well as the direct SSI at normal incidence, for various durations ranging from 1 min to 1 month.

Since its inception as a pre-operational service, McClear has been increasingly used by academics and practitioners. A lot of attention is paid to the validation of the estimates provided by McClear. The goal is to better establish the domain of validity of McClear, its qualities and drawbacks, and to bring transparency and confidence in the use of this operational service.

McClear has been previously validated with respect to 1 min measurements of global and direct SSI on horizontal surface from the Baseline Surface Radiation Network (BSRN) collected from 11 sites located throughout six continents (Lefèvre et al., 2013). The relative root mean square error (RMSE) for global SSI and direct SSI depends on the station and ranges respectively between 3 and 5% of the mean of the measurements for the station, and between 5 and 10%.

This article aims at contributing further to the validation of the McClear model. It focuses on desert conditions encountered in Israel where three close stations measure the global, diffuse and direct SSI. This density of stations permits to study the variability of the performances of McClear in this climate homogeneous area.

2 Measurements and McClear estimates

Measurements of the global *G* and diffuse *D* SSI and of the beam irradiation received at normal incidence B_N were collected from three stations (Fig. 1 and Table 1) from the Israel Meteorological Service (IMS), the BSRN network and an undisclosed company for the period 2006–2011. The direct SSI *B* on horizontal surface is computed from the difference *G*–*D*. Measurements are integrated over 10 min at Beer Sheva and Yotvata and 1 min at Sede Boqer which belongs to the BSRN network. 1 min measurements at Sede Boqer were averaged over 10 min to match the sampling rate of the two other stations. The solar zenith angle θ_S corresponding to each measurement is computed with the SG2 al-



Figure 1. Map of the three stations. The red line is 100 km in length.

gorithm (Blanc and Wald, 2012). After applying the quality check procedures of Roesch et al. (2011), only the measurements were kept which pass the filters proposed by Lefèvre et al. (2013) to retain reliable clear-sky instants. Finally, only are kept clear-sky instants for which measurements are valid for the three stations simultaneously. The number of samples is 19 849 for each station. This last constraint was imposed in order to be able to compare correlation coefficients computed for data sets, whether measurements or estimates, for two stations.

The three stations are fairly close to each other (Fig. 1). Beer Sheva is 40 km north of Sede Boqer and Yotvata is 120 km south of Sede Boqer.

McClear estimates of *G*, *D*, *B* and *B*_N for 10 min duration were obtained from the SoDa web site (www.soda-pro.com) for these same instants and for each location. It may be of interest here to underline that the McClear model computes *G* and *B*_N, then *B*, and that *D* is deduced from *G*–*B*. Also provided were the corresponding time-series of the irradiance at the top of atmosphere on both horizontal and normal surfaces: E_0 and E_{0N} . The clearness index *KT* and the direct clearness index *KT*_{BN} were computed for both measurements and McClear estimates using the following formula:

$$KT = G/E_0 \tag{1}$$

$$KT_{B_{\rm N}} = B_{\rm N}/E_{0{\rm N}}.$$

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Table 2. Comparison between clear-sky global G and diffuse D SSI measured by ground stations and estimated by McClear. Units in $W m^{-2}$.

		G			D	
	BEE	SBO	YOT	BEE	SBO	YOT
Mean observed SSI	810	838	825	124	114	137
Bias	19	2	32	60	69	48
Relative bias (%)	2	0	4	48	60	35
RMSE	32	30	41	66	74	55
Relative RMSE (%)	4	4	5	53	65	40
Squared correlation coefficient	0.977	0.976	0.980	0.594	0.627	0.633



Figure 2. 2-D histogram of measurements (horizontal axis) and McClear estimates (vertical axis) of *G* for Sede Boqer. The colour represents the frequency of each pair.

3 Results

Following the ISO (International Organization for Standardization) standard (1995), the deviations were computed by subtracting measurements for each instant from the McClear estimates and they were summarized by the bias, the root mean square error (RMSE), and the squared correlation coefficient, also known as the coefficient of determination (R^2). Relative values are expressed with respect to the mean observed value. The validations of KT and KT_{B_N} are also included, as they are stricter measures of the performance of a model with respect to the optical state of the atmosphere.

The 2-D histograms of measured and estimated values are presented for Sede Boqer (Figs. 2 and 3). Red, respectively dark blue, dots correspond to regions with great, respectively very low, densities of samples. The plots also present the number of samples, the mean reference value, the bias, the RMSE, the correlation coefficient (CC) and the 1 : 1 line (y = x). One may see in Fig. 2 that the points are mostly

Sede Boger: beam normal clear sky irradiation (W m⁻²)



Figure 3. 2-D histogram of measurements (horizontal axis) and McClear estimates (vertical axis) of B_N for Sede Boqer. The colour represents the frequency of each pair.

aligned with the 1:1 line with a very limited scattering. The bias and RMSE are respectively 2 and 30 W m⁻². The squared correlation coefficient is very large: 0.976, meaning that the temporal changes in *G* are well reproduced by Mc-Clear. The points in Fig. 3 for B_N are less aligned with the 1:1 line. The absolute value of the bias and RMSE are much larger: -68 and 83 W m⁻². The squared correlation coefficient is 0.610 and a large amount of changes in B_N is unexplained by McClear.

Tables 2–4 present the results of the comparison for respectively *G*, *D*, *B*, *B*_N, *KT* and *KT*_{*B*_N}. The means of *G* (Table 2, approximately 830 W m⁻²), *B*_N (Table 3, approximately 840 W m⁻²) and clearness indices (Table 4, 0.75 and 0.64) are large which means that the atmosphere is very often clear and not turbid. Yotvata experiences less B_N – and a lower KT_{B_N} – though it is the southernmost site. It is located 40 km north of the Red Sea in the Negev desert and may be under maritime influence and dust episodes. **Table 3.** Comparison between clear-sky beam SSI *B* and beam at normal incidence B_N measured by ground stations and estimated by McClear. Units in W m⁻².

		В			$B_{\rm N}$	
	BEE	SBO	YOT	BEE	SBO	YOT
Mean observed SSI	686	724	688	841	878	809
Bias	-41	-66	-16	-46	-68	13
Relative bias (%)	-6	-9	-2	-6	-8	2
RMSE	59	80	46	69	83	53
Relative RMSE (%)	9	11	7	8	10	7
Squared correlation coefficient	0.931	0.937	0.929	0.576	0.610	0.603

Table 4. Comparison between clear-sky clearness indices KT and KT_{B_N} from ground stations and estimated by McClear.

		KT			KT_{B_N}	
	BEE	SBO	YOT	BEE	SBO	YOT
Mean observed index	0.74	0.76	0.74	0.62	0.66	0.62
Bias	0.02	0.01	0.03	-0.04	-0.06	-0.01
Relative bias (%)	3	1	5	-6	-9	-2
RMSE	0.03	0.03	0.05	0.05	0.07	0.04
Relative RMSE (%)	5	4	6	8	11	7
Squared correlation coefficient	0.524	0.471	0.474	0.625	0.641	0.578

The bias for G is low for Sede Boqer: 2 W m^{-2} , and is larger for the other sites: 19 and 32 Wm^{-2} , i.e. 2 and 4% of the mean observed G. The RMSE ranges from 30 to $41 \,\mathrm{W \,m^{-2}}$ (4%) and the squared correlation coefficient is greater than 0.976 (Table 2). The influence of $\theta_{\rm S}$ on the SSI creates de facto a correlation between measurements and estimates in clear-sky conditions as $\theta_{\rm S}$ and E_0 can be accurately estimated. The influence of $\theta_{\rm S}$ on KT is much less pronounced and the squared correlation coefficient denotes the ability of McClear to reproduce the optical state of the atmosphere. It ranges between 0.471 and 0.524 (Table 4) and is low. A majority of changes in KT is not reproduced by Mc-Clear and improvements should be brought on the McClear model and on the quality of its inputs. The bias and RMSE for *KT* are similar in relative value to those for *G* (Table 4). Expectedly, the results for Sede Boger are fully in line with the bias, RMSE and squared correlation coefficient for both G and KT reported by Lefèvre et al. (2013) for this station though for 1 min SSI: 7 and 30 W m⁻², 0.982, and 0.01, 0.03 and 0.581.

The estimates of *D* by McClear are inaccurate (Table 2). There is an overestimation ranging between 48 and 69 W m⁻² (35 to 60 % of the mean of *D*). The RMSE ranges between 55 and 74 W m⁻² (40 to 65 %). The squared correlation coefficient is comprised between 0.594 and 0.633; a large amount of changes in *D* is unexplained by McClear.

An underestimation is observed for *B* and B_N (Table 3), except Yotvata for B_N . The bias for *B*, respectively B_N , is comprised between -66 and -16 W m⁻², i.e. between -9

and -2% of the mean *B*, and between -68 and $+13 \text{ W m}^{-2}$ (-8 and 2% of the mean *B*_N). The RMSE ranges from 46 (7%) to 80 W m⁻² (11%) for *B*, and from 53 (7%) to 83 W m⁻² (10%) for *B*_N. The bias and RMSE for *KT*_{*B*_N} are similar in relative value to those for *B* and *B*_N (Table 4). The squared correlation coefficient for *B*_N and *KT*_{*B*_N} is close to 0.6; a large amount of changes in *B*_N or *KT*_{*B*_N} is unexplained by McClear. The squared correlation coefficient for *B* is much larger and close to 0.93 because of the influence of $\theta_{\rm S}$ on the correlation and the accuracy of its estimate.

An additional comparison was performed that dealt with the ability of McClear to reproduce spatial variability. The correlation coefficient between time-series of measurements, respectively McClear estimates, was computed for each pair of stations for *G* and B_N (Table 5). It is observed (upper right part of the correlation matrix) that the measurements are very much correlated for *G* (greater than 0.99), which can be explained by the fact that only clear-sky measurements are dealt with. The correlation coefficient is less for B_N , especially between Yotvata and the two others for which it is respectively 0.630 and 0.728. This is in agreement with the remoteness of Yotvata compared to the two others and the above remark on its climate.

The closer the correlation coefficients of the lower part of the matrix to those of the upper part, the more accurately Mc-Clear depicts the variability in space. The correlation coefficients for *G* are almost identical for the measurements and McClear meaning that the actual SSI field is well reproduced by McClear. This is not the case for B_N for which discrep-

Table 5. Correlation matrix between stations for measurements (upper right part of the matrix, in bold) and for McClear (lower left part, in italic) for G and B_N .

G	BEE	SBO	YOT	B _N	BEE	SBO	YOT
BEE	1	0.995	0.991	BEE	1	0.884	0.630
SBO	0.999	1	0.994	SBO	0.990	1	0.728
YOT	0.994	0.996	1	YOT	0.846	<i>0.908</i>	1

ancies may be observed. There is an overestimation of the correlation by McClear which can be attributed to the correlation of its inputs due to the coarse spatial and temporal resolutions. CAMS products on aerosols and total content in water vapour and ozone are available every 3 h. The spatial resolution is 1.125° , i.e. approx. 120 km along a longitude, for the aerosol properties. This is the same resolution for the total column content of ozone and water vapour before 2014 after which it became 0.8° . The $B_{\rm N}$ field estimated by McClear will be smoother than the actual field. Note that the ranking of the correlation coefficients is the same for both the measurements and McClear; the local extrema are respected though the intensity of the variation is decreased.

4 Discussion and conclusion

Like reported in other similar studies, the statistical quantities reported here vary with the period of analysis. A given quantity may change noticeably from one year to another. For example, the bias in B_N at Sede Boqer varies from -63 to -75 W m⁻² if years are considered separately. This indicates that care must be taken in the analysis of these quantities.

The quantities vary with the month. Trends are more or less marked. There is a tendency for lowest bias – in absolute value – and lowest RMSE in the period May–August. There is a tendency for the bias and the RMSE to increase with θ_S , yielding an increase – in absolute value – of the relative bias and RMSE as the mean *G* and *B*_N decrease as θ_S increases. Nevertheless, the changes are limited.

Eissa et al. (2015a, b) have performed similar studies but for respectively Egypt and the United Arab Emirates. Similarly to this study, Aswan and the UAE sites exhibit underestimation of $B_{\rm N}$. This underestimation is more pronounced for Beer Sheva and Sede Boger. On the contrary, Yotvata exhibits an overestimation of $13 \text{ W} \text{ m}^{-2}$. The comparison of these different studies shows that the overall picture of the possible causes of the discrepancies between measurements and Mc-Clear estimates is still unclear. The underestimation in $B_{\rm N}$ may be partly caused by overestimation of the aerosol optical depth (AOD). Through comparisons between the AODs measured by AERONET and estimated in CAMS for desert areas in Egypt and UAE, Eissa et al. (2015b) and Oumbe et al. (2012) concluded that one main source of the errors in McClear originates from the CAMS AOD. Therefore, more accurate inputs to McClear would improve its estimates. For example, Oumbe et al. (2015) have shown that a local empirical correction of the CAMS AOD drastically decreases the bias in the United Arab Emirates.

As for *G*, if one looks at the results of Eissa et al. (2015a) for Aswan in Egypt – which is located in a desert far from the Cairo megapole, – one would observe that the large overestimation of *G* by McClear over Aswan: 33 W m^{-2} , is similar to that observed at Yotvata in this study. Yotvata exhibits the greatest bias of the three sites. The bias at Sede Boqer is 2 W m^{-2} , expectedly similar to that of 7 W m^{-2} reported by Lefèvre et al. (2013) for the same site though for 1 min summarization. The bias for the more turbid sites in the UAE ranges from -5 to 10 W m^{-2} .

Estimates in G and D – and hence the statistical performances - are sensitive to the type of aerosols that is estimated by the means of the empirical algorithm presented in Lefèvre et al. (2013) applied to the partial aerosol optical depths delivered by CAMS. It is found that in Beer Sheva and Sede Boger – which are close compared to the size of the CAMS cell, - the most frequent aerosol type is "continental polluted", then "maritime polluted" and finally "desert". The same types are found for Yotvata but "desert" is most frequent than "maritime polluted". An error may arise if the wrong type is selected. Figure 1 in Lefèvre et al. (2013) displays a specific case of daily profile of G in Carpentras (France) with a dramatic change by $30 \,\mathrm{W}\,\mathrm{m}^{-2}$ (approx. 3%) due to an error in the empirical algorithm. In other cases reported in Eissa et al. (2015a) an overestimation of the fine, strongly scattering pollution particles associated with an underestimation of the coarse, less scattering, mineral dust particles would affect G and D. It should be added that the coarse spatial and temporal resolutions of the CAMS data on aerosols make it difficult to capture the exact atmospheric effects on the incident solar radiation over a specific site. Other causes of uncertainty are the uncertainties in the OPAC model used in McClear (Lefèvre et al., 2013). Zieger et al. (2010) showed noticeable changes in single scattering albedo with relative humidity for the OPAC "continental polluted" and "maritime polluted" types. If relative humidity is assumed too large, then the single scattering albedo is overestimated, yielding an overestimation in D. This may explain the difference between the two sites Beer Sheva and Sede Boger and the southern one Yotvata where "desert" is more frequent. Simulations performed with the radiative transfer model libRadtran have shown that in case of intense dust storms, i.e. heavy load in dust particles, the single scattering albedo in OPAC "desert" type underestimates that observed in AERONET measurements, which yields an underestimation in D. This is not observed in cases of low or medium loads in dust. This adds to the complexity as intense dust storms may also be observed in the northern sites.

Performances are still far from WMO standards: bias less than 3 W m^{-2} and 95 % of the deviations less than 20 W m^{-2} . Uncertainties in aerosol properties from CAMS are still too

large, and more efforts are necessary for a better modelling of the aerosols.

Despite the identified drawbacks and paths for improvements, this validation of the McClear service for the desert conditions in Israel reveals satisfactory results. The comparisons between the McClear estimates and measurements of global horizontal and direct normal irradiances for 3 stations show that a large correlation is attained showing the ability of McClear to capture the temporal and spatial variability of the irradiance field. The performances are similar for the three sites for the global irradiance and for the DNI to a lesser extent, demonstrating the robustness of the CAMS McClear service.

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